

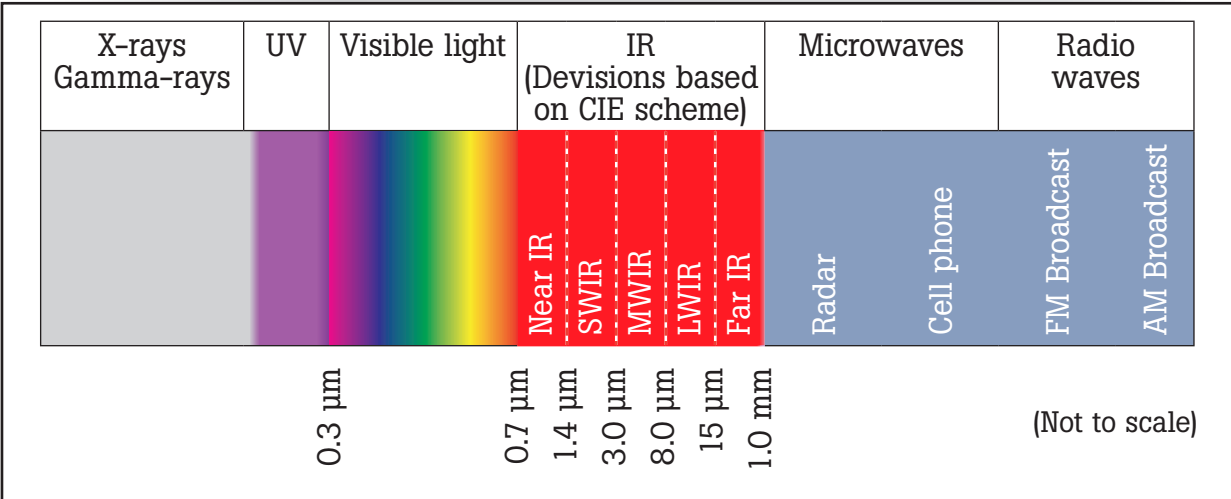
# Quick reference to infrared thermometer calibration

FLUKE®

Calibration

## Infrared theory

### IR and the electro-magnetic spectrum



### Blackbody theory

#### Planck's Law

Describes the amount of spectral radiance emitted by a perfect blackbody at a given temperature, wavelength and solid angle.

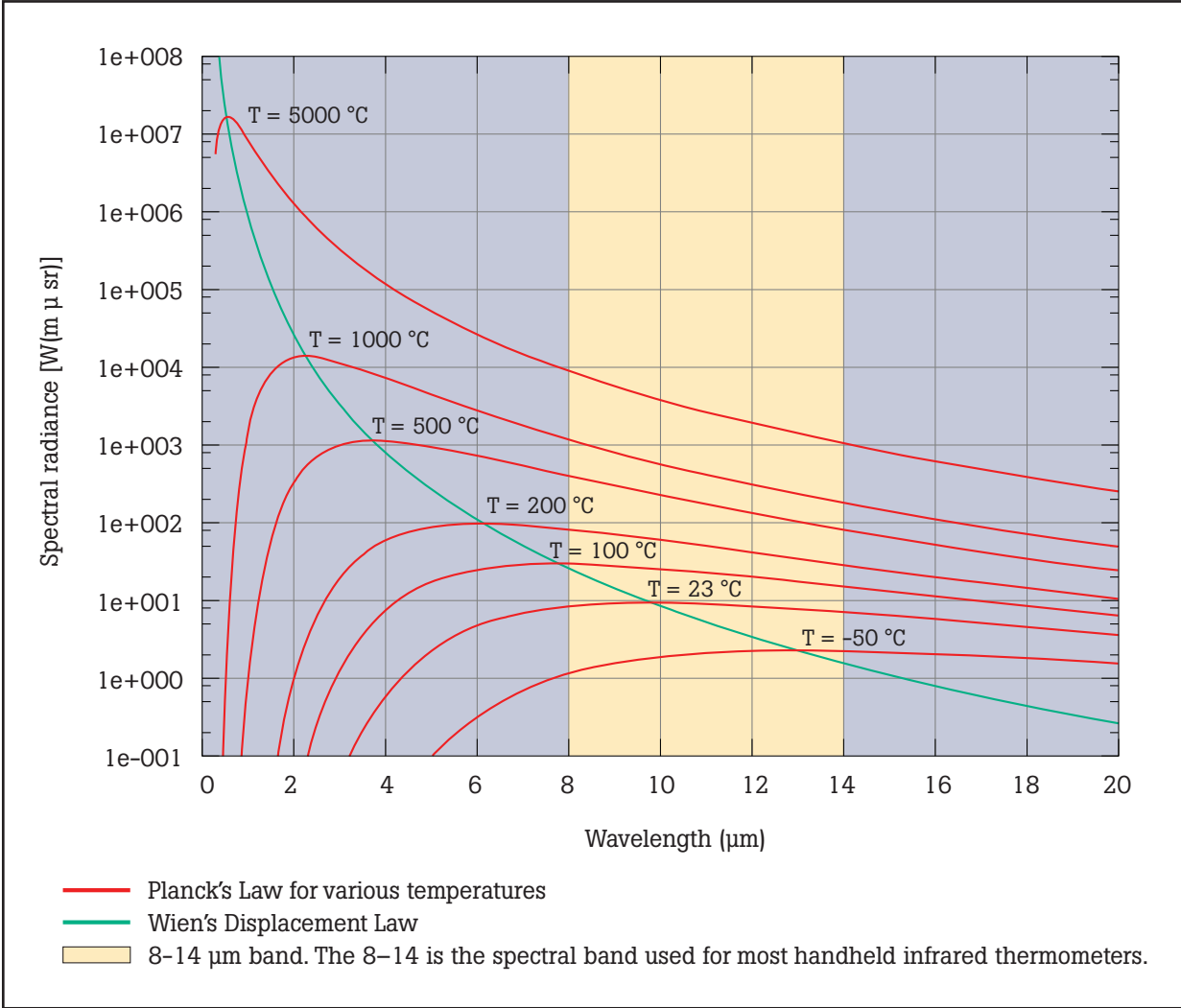
$$L(\lambda, T) = \frac{C_{1L}}{\lambda^5 \left[ \exp\left(\frac{C_2}{\lambda T}\right) - 1 \right]}$$

#### Wien's Displacement Law

Shows the peak wavelength for a given temperature (as predicted by Planck's Law).

$$\lambda_{\text{PEAK}} T = c_3$$

#### Planck's Law and Wien's Displacement Law



#### Stefan-Boltzmann Law

Shows the total irradiance in the entire electro-magnetic spectrum.

$$M = \sigma T^4 = \pi \int_0^\infty \frac{C_{1L}}{\lambda^5 \left[ \exp\left(\frac{C_2}{\lambda T}\right) - 1 \right]} d\lambda$$

#### Physical constants

Name	Symbol	Value
First radiation constant for spectral radiance	$c_{1L}$	$1.191\,042\,759\,e^8\,W\,\mu m^2\,sr^{-1}$
Second radiation constant	$c_2$	$1.438\,7752\,e^4\,\mu m\,K$
Wien Wavelength Displacement Law constant (Third radiation constant)	$c_3$	$2.897\,7685\,e^3\,\mu m\,K$
Stefan-Boltzmann constant	$\sigma$	$5.670\,400\,e^{-8}\,Wm^{-2}\,K^{-4}$

#### Sakuma Hattori Equation (Planckian Form)

The Sakuma-Hattori Equation is used to approximate Planck's Law for instruments with non-finite bandwidth. It can be used for a determination of temperature from radiance, or it can be used for determination of uncertainties using generalized constants.

$$S = \frac{C}{\exp\left(\frac{C_2}{AT + B}\right) - 1}$$

$$T = \frac{C_2}{A \ln\left(\frac{C}{S}\right) + 1} - \frac{B}{A}$$

Generalized constants for 8-14 μm Band:

A = 9.364 μm  
B = 178.4 μm K  
C = 1.000 (unitless)

$$\frac{\partial S}{\partial T} = [S(T)]^2 \frac{Ac_2}{C(AT + B)^2} \exp\left(\frac{c_2}{AT + B}\right)$$

#### Use Kelvins

When doing calculations with any of the equations above, be sure to do the calculations using kelvins, not degrees Celsius.

#### Kirchhof's Law and Emissivity

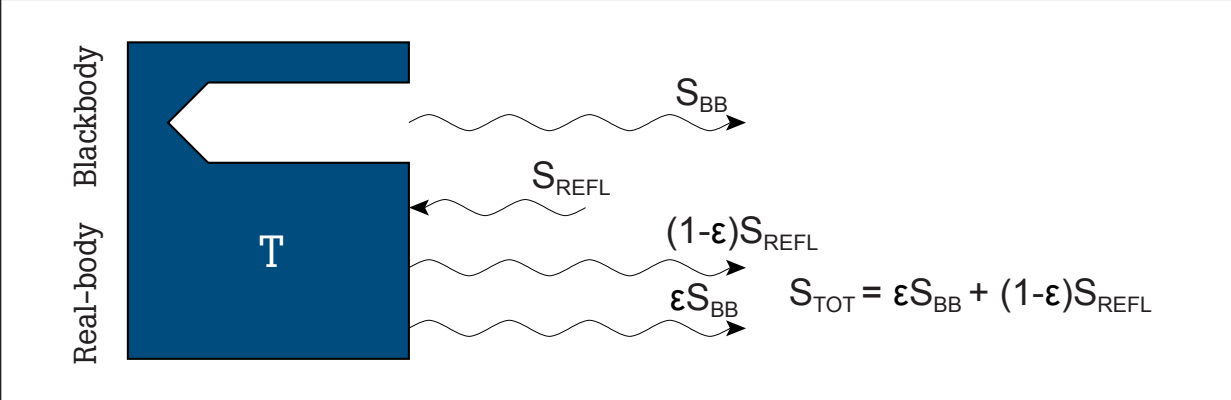
For energy incident on a surface, the ratios of the energy reflected, transmitted and emitted sum to unity. When a surface is opaque, the emitted and reflected energy sum to unity. The emissivity is the ratio a surface of temperature (T) emits when compared to a perfect blackbody of the same temperature.

$$1 = \epsilon + \rho + \tau$$

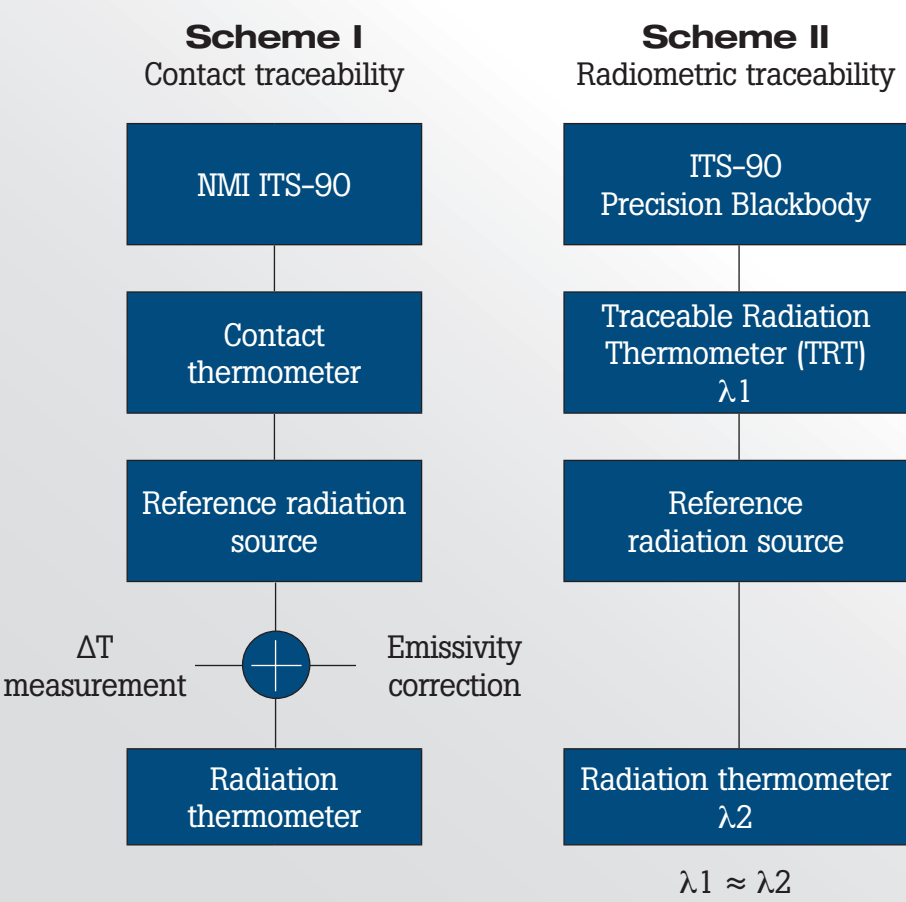
For opaque surfaces:

$$\tau = 0 \rightarrow 1 = \epsilon + \rho$$

Every object has an emissivity less than unity. The energy exiting an opaque surface is a combination of reflected and emitted energy. It is important to note that emissivity is not necessarily constant over all wavelengths.



### Calibration schemes



### Uncertainty budget structure

#### Blackbody

- Calibration temperature (VTBB)
- Impurities (FPBB)
- Plateau identification (FPBB)
- Blackbody emissivity, isothermal
- Blackbody emissivity, non-isothermal (VTBB)
- Reflected ambient radiation
- Cavity bottom heat exchange
- Convection (VTBB)
- Cavity bottom uniformity (VTBB)
- Ambient conditions (VTBB)
- VTBB: Variable Blackbody scheme only
- FPBB: Fixed Point Blackbody scheme only

#### Radiation thermometer

- Size-of-source effect
- Non-linearity
- Reference temperature
- Ambient temperature
- Atmospheric absorption
- Gain ratios
- Noise

#### Use

- Interpolation error
- Drift
- Unknown temperature

(based on BIPM CCT-WG5 on Radiation Thermometry, Uncertainty Budgets for Calibration of Radiation Thermometers below the Silver Point)

### Measurement equation

#### Generalized measurement equation

The true temperature of the surface being measured.

$$S(T_{\text{MEAS}}) = S(T_{\text{SURF}}) + \frac{(1 - \epsilon_{\text{IRT}})}{\epsilon_{\text{IRT}}} [S(T_{\text{REFL}}) - S(T_{\text{DET}})] + \frac{(\epsilon_{\text{SURF}} - \epsilon_{\text{IRT}})}{\epsilon_{\text{IRT}}} [S(T_{\text{SURF}}) - S(T_{\text{REFL}})]$$

Account for the effects of uncertainty in reflected temperature.

Account for the effects of uncertainty in emissivity.

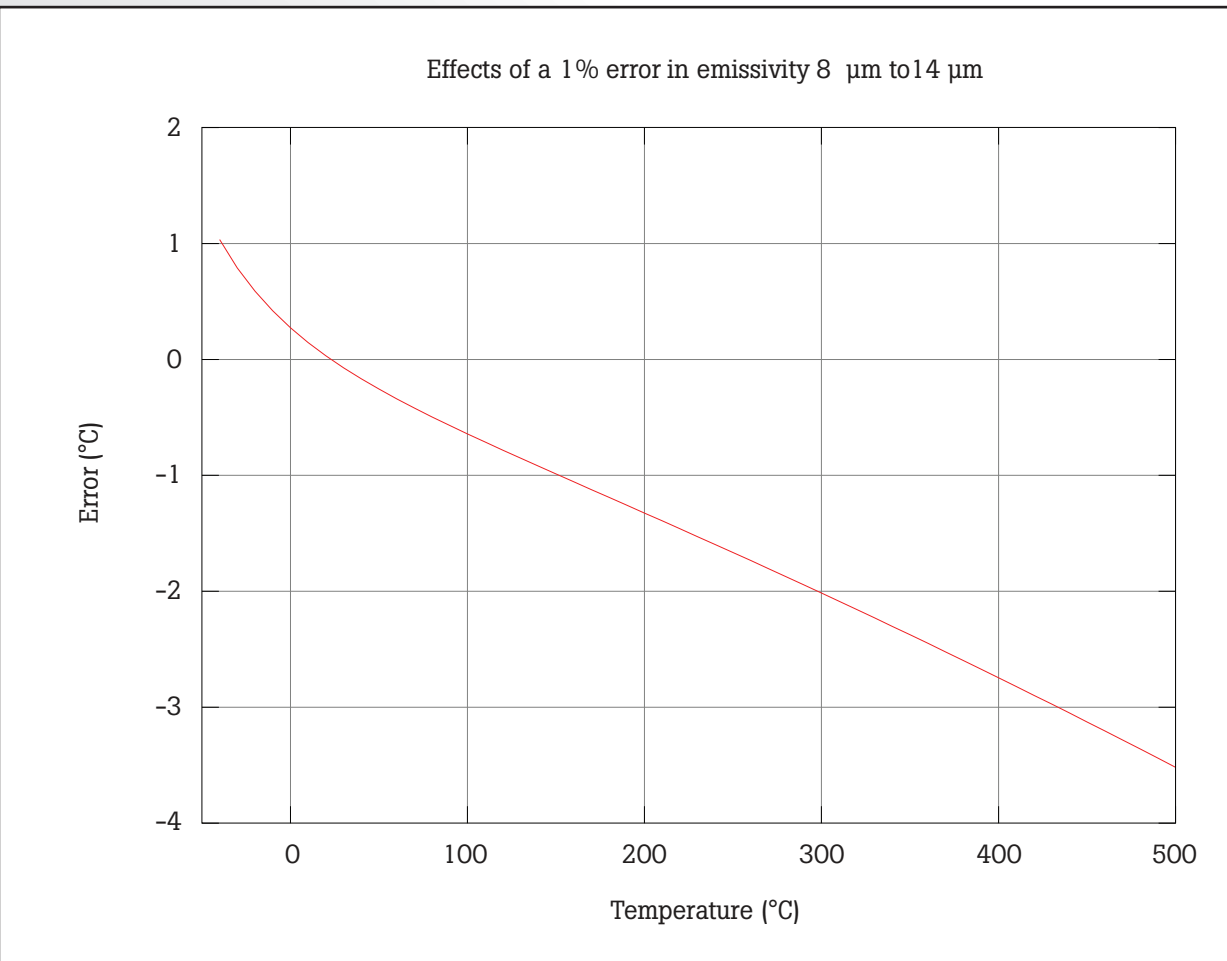
To calculate the error based on any of these elements, simply perform the following calculation: error = S(T<sub>MEAS</sub>) - S(T<sub>SURF</sub>)

#### Effects of emissivity uncertainty

To calculate the effects of emissivity uncertainty, ASTM E2758 provides the equation to the right. Any S(T) involves a calculation using the Sakuma-Hattori Equation.

$$\text{error} = \frac{(\epsilon_{\text{SURF}} - \epsilon_{\text{IRT}})}{\epsilon_{\text{IRT}}} [S(T_{\text{SURF}}) - S(T_{\text{REFL}})]$$

For the various spectral bands, the following graph shows the results error based on the following conditions.



#### Effects of reflected temperature

$$\text{error} = \frac{(1 - \epsilon_{\text{IRT}})}{\epsilon_{\text{IRT}}} [S(T_{\text{REFL}}) - S(T_{\text{DET}})]$$

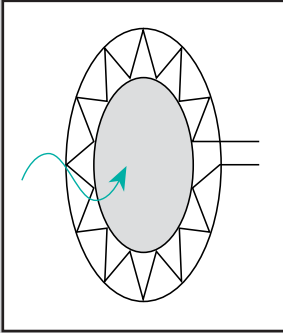
## Infrared thermometers

### What is an infrared thermometer?

Infrared thermometers are a subset of radiation thermometers. These devices measure infrared radiation and display a temperature based on the radiation measured by the infrared thermometer and the emissivity setting of the infrared thermometer. The term infrared thermometer generally refers to handheld devices with a thermopile detector. Some other names used for infrared thermometers are IR guns, point and shoot thermometers, spot pyrometers, laser thermometers.

### Thermopile detectors

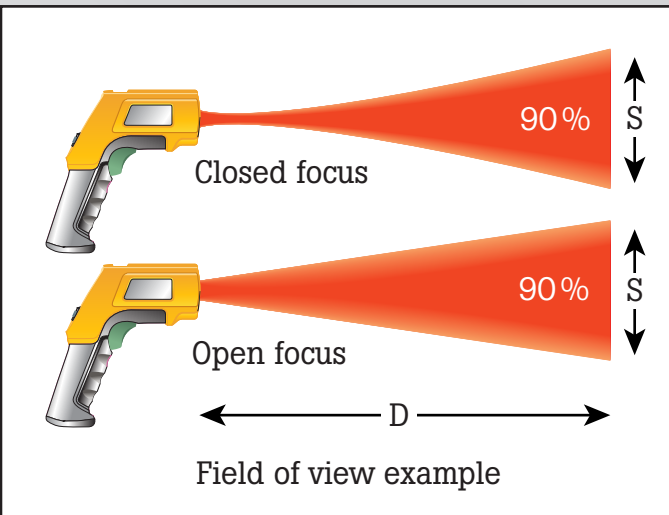
Essentially, a thermopile is a number of thermocouple junctions joined in series. The radiation strikes a drum-like disk, causing it to heat. A ring forms a surface for the reference temperature. The voltage of the thermopile is measured and a temperature is assigned to that voltage, similar to what is done with thermocouple measurement.



### Infrared thermometer optics

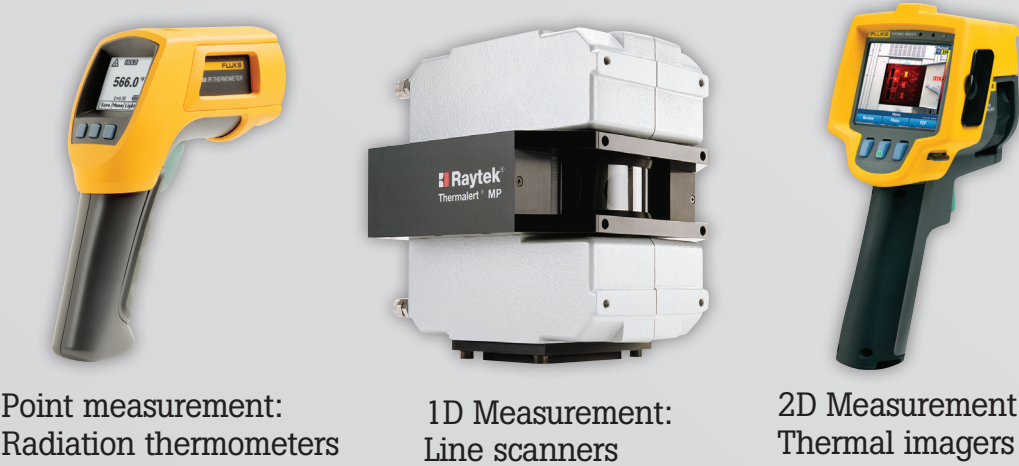
#### Size-of-source effect, field of view and D:S ratio

Size-of-source, field-of-view and distance to size ratio refer to optical characteristics of the IR thermometer. In the case of field of view and distance to size ratio, both specifications refer to a diameter, a measuring distance and a percentage of energy within the diameter. This means that even though a radiation thermometer has a given source diameter, there is still some measurement coming from outside that diameter. This is called scatter.



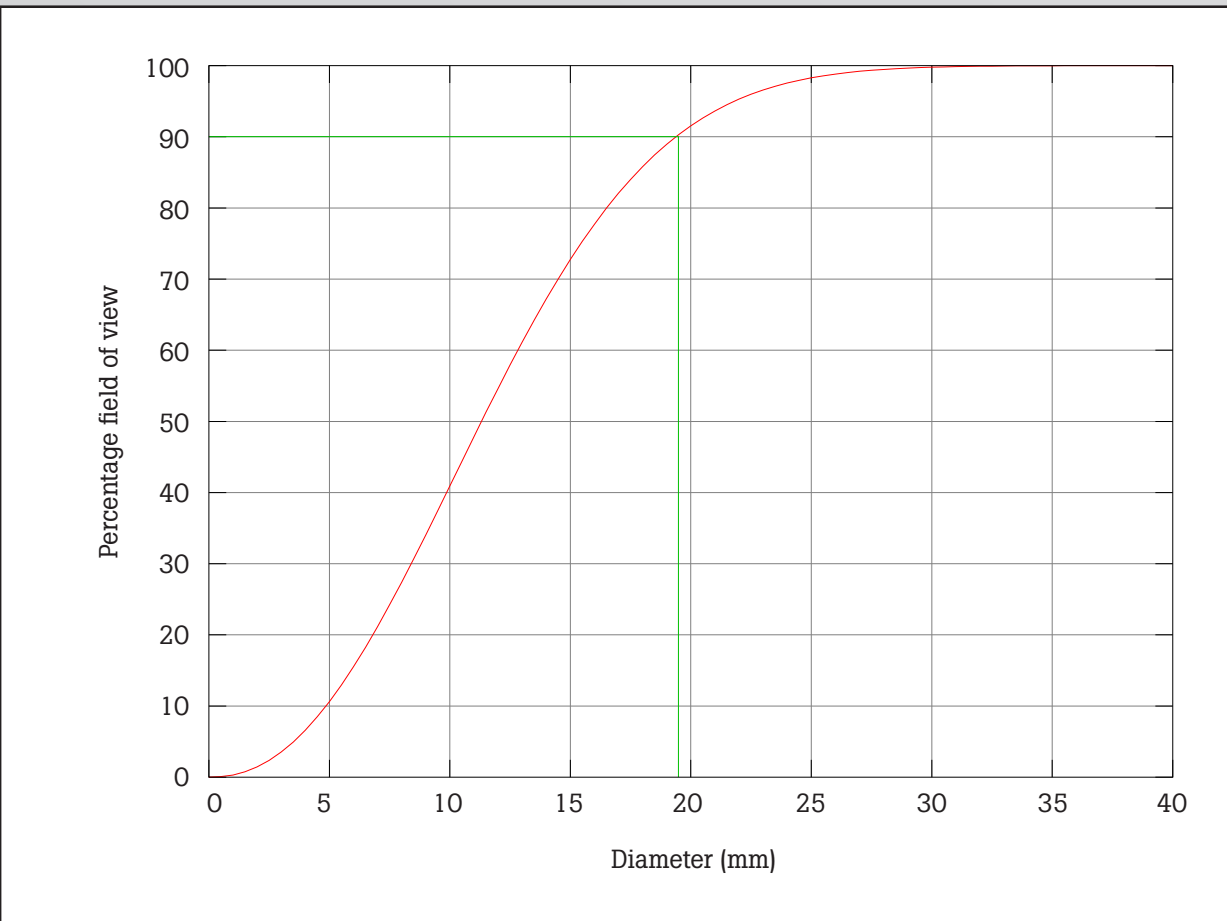
In the example to the right, a radiation thermometer's field of view has been measured at various diameters at a fixed measuring distance. In this example, the radiation thermometer measured 90% of the target's energy within a diameter of 19 mm. If the measuring distance is 400 mm, then the field of view would be specified as: 19 mm diameter (90%), measuring distance: 400 mm. Source: IEC/TS 62492-1

### Classification of instruments

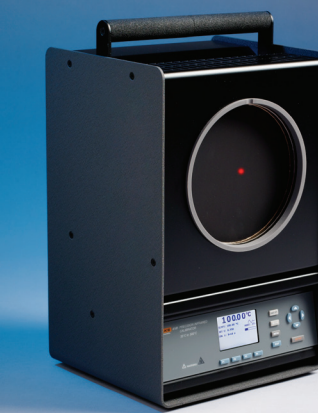


### Infrared thermometer models

Model	Measuring range	D:S
Fluke 561	-40 °C to 550 °C (-40 °F to 1022 °F)	12:1
Fluke 566	-40 °C to 650 °C (-40 °F to 1202 °F)	30:1
Fluke 568	-40 °C to 800 °C (-40 °F to 1472 °F)	50:1
Fluke 572	-30 °C to 900 °C (-40 °F to 1652 °F)	50:1 or 60:1
Fluke 574	-30 °C to 900 °C (-40 °F to 1652 °F)	50:1 or 60:1



### The Fluke Calibration 4180 and 4181 Precision Infrared Calibrators



- Radiometric calibration to account for emissivity
- Large enough target to reproduce calibration geometry
- Thermometer emissivity setting compensation